# Analyzing Hyperspectral BRDF Data of a Grass Lawn and Watercress Surface Using an Empirical Model

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### ABSTRACT

This paper presents results from analyzing hyperspectral BRDF data of grass lawn and watercress. The intensity of the hotspot as a function of wavelength is determined from fitting an empirical (or rather *phenomenological*) model to the data. The model consists of a lambertian, bowlshape, hotspot and forward scattering component. An analytical, theoretically derived relation between the hotspot intensity and the lambertian component is given. The intensity of the multiple scattered radiation obtained from this relation can be explained qualitatively with the surface structures of the samples.

## INTRODUCTION

The Bidirectional Reflectance Distribution Function (BRDF) describes the dependence of surface reflectance on incidence and viewing angles as a function of wavelength. With hyperspectral BRDF data, the dependence of the BRDF on the 'brightness' of the surface at the respective wavelength can be studied effectively on a vegetation canopy, since the spectral albedo of vegetation varies considerably (e. g. from 0.026 to 0.43 between 500 nm and 1000 nm for a grass sample). The spectral signature affects the shape of the BRDF, because multiple scattering increases dramatically at high albedos. This paper focuses on the decrease of the hotspot relative to the albedo as multiple scattering increases.

## DESCRIPTION OF DATA SETS

Data from an erectophile grass (*Lolium perenne*) and a planophile watercress (*Lepidium sativum*) surface were acquired under controlled laboratory conditions at the European Goniometric Facility (EGO) of the Joint Research Center in Ispra/Italy [1]. The angular resolution of the BRDF data is  $5^{\circ}$  and  $15^{\circ}$  in zenith and azimuth direction, respectively. The data is available from our website. In addition, a field data set of the same grass species is analyzed, measured with the FIGOS field goniometer with a resolution of  $15^{\circ}$  and  $30^{\circ}$  in zenith and azimuth, respectively [2]. All data were taken under a source zenith angle of  $35^{\circ}$  with zenith angles of reflection  $\theta_r$  ranging from  $0^{\circ}$  to  $75^{\circ}$  using a GER-3700 spectrora-

diometer with 323 channels between 500 nm and 1000 nm.

#### EMPIRICAL BRDF MODEL

Four basic BRDF components are identified and fitted to the data with an empirical function: 1) hotspot 2) forward scattering 3) bowlshape and 4) lambertian component.

The forward scattering component is adopted from a specular term successfully used in similar studies [3] (see [4] for a detailed explanation of the specular term). The bowl shape is modeled as a simple linear function of the zenith angles.

To model the hotspot, we chose a simple exponential function  $e^{-b_2 \cdot g}$  which has a shape very similar to a function proposed by B. Hapke [5] (g is the relative angle between viewing and illumination direction). Our function has the advantage of approaching zero in the forward scattering direction much faster than Hapke's function, which makes it easier to separate the hotspot from the forward scattering term in the inversion step. Since the decrease of reflectance from the hotspot towards larger viewing zenith angles is smaller than the decrease towards nadir, we multiplied the hotspot term with an exponential function  $e^{b_1 \cdot (\theta_i \cdot \theta_r)^2}$  also used in the forward scattering component.

The resulting BRF model has the following form:

$$BRF = \pi \cdot BRDF = a_0 + a_1 \cdot (\theta_i + \theta_r) + a_2 \cdot e^{b_1 \cdot (\theta_i \cdot \theta_r)^2} \cdot e^{-b_2 \cdot g} + a_3 \cdot e^{b_3 \cdot (\theta_i \cdot \theta_r)^2} \cdot e^{b_4 \cdot \psi^2}$$
(1)

where  $\psi$  is the relative angle to the specular direction, see [4]. The coefficients  $a_i$  depend on wavelength, the coefficients  $b_i$  are fixed for each sample. The underlying assumption is that the shape of the hotspot and the forward scattering term are primarily determined by the geometry of the canopy, whereas the intensity of these two terms can vary with wavelength due to the change of reflectance with wavelength.

The coefficients have the following physical meaning:

- $a_0$  = lambertian component (diffuse scattering)
- $a_1$  = intensity of the bowlshape
- $a_2 =$  intensity of the hotspot
- $a_3$  = intensity of the specular peak

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 $b_2$  = width of the hotspot

For an explanation of  $b_1$ ,  $b_3$  and  $b_4$  see [4].

Based on the assumption that the surface consists of surface patches of one reflectance only (so the background does not influence the BRDF), we derived a *physical* (not empirical) relation between  $a_0$  and  $a_2$ :

$$a_{2} = \frac{e^{-b_{1}\theta_{i}^{*}}}{2\pi k} \cdot \left(-K_{1} + \sqrt{K_{1}^{2} - 4\pi kK_{0}}\right) \quad \text{with} \qquad (2)$$
  
$$K_{0} = a_{0}(x - c) + a_{0}^{2}\pi k(1 - c), \quad K_{1} = a_{0}(2\pi k - c) + x$$

where  $x \in [0, 1]$  is the areal proportion of shadowed surface patches at angles where the BRDF equals  $a_0$  (diffuse limit),  $k \in [0, 1]$  is the probability that scattered light will hit another surface patch, and c is a geometry factor (close to 1). x, k and c are wavelength independent quantities. The derivation of (2) and the assumptions made will be presented in a later publication.

## MODEL INVERSION

We used a numerical routine from the software package 'PV-WAVE' based on the least-squares-deviation to fit the parameters of equation 1 to the BRDF data. Allowing all parameters to float freely did not yield reasonable results. So we adopted a step by step procedure.

First, we fitted parameter  $a_1$  for the bowlshape component for each wavelength taking into account only BRDF measurements in the cross principle plane (relative azimuth  $\varphi = 90^{\circ}$ ) where the effects of the hotspot and the forward scattering term are assumed to be smallest. Then, we estimated parameters  $b_i$  (the shape-determining parameters of the hotspot and forward scattering term) from BRDF measurements in the principal plane ( $\varphi = 0^{\circ}/180^{\circ}$ , respectively), taking into account the effect of parameter  $a_1$ . For the forward scattering term, some freedom remains in the choice of  $b_3$  and  $b_4$ . This term was designed to fit to a specular peak with a distinct maximum, but our data only shows a rise and no decline (if there is a maximum, it must be at  $\theta_r > 75^\circ$ ). The parameters  $b_1$ and  $b_2$  could be determined well, because the hotspot peak (at  $\theta_i = \theta_r = 35^\circ$ ) is well within the range of measured angles  $(\theta_r = [0^\circ ... 75^\circ])$ . For the field measurement of grass, the same coefficients  $b_i$  were used as for the laboratory measurement of grass because the angular resolution of the field measurement was much coarser.

In the final step, we fitted the parameters  $a_0$ ,  $a_2$  and  $a_3$  to all data points.

## **RESULTS AND DISCUSSION**

The agreement between measurements and fitted function (1) is satisfying, the mean relative deviation is only a few %, see table 1. The assumption that the shape of the hotspot does not change with wavelength was confirmed clearly for the grass laboratory sample. The coefficients  $b_i$  are given in table 1.  $b_1$  equals 1.5 for all samples.  $b_2$  for cress is greater than  $b_2$  for grass, this means that cress has a sharper hotspot. The coefficients  $a_i$  are shown in Fig. 1. All 3 samples show the rising edge characteristic for vegetation at about 700 nm, see plots

A-C (neglecting BRDF effects, the coefficient  $a_0$  is approximately proportional to the albedo).

The bowlshape component described by  $a_1$  for grass is very different from cress (plots D-F).  $a_1$  varies only by a factor of about 2 for cress (cress has almost no bowlshape at NIR), whereas for grass,  $a_1$  goes to zero for small  $a_0$ . Thus,  $a_1$  for grass is roughly proportional to  $a_0$  for  $a_0 < 0.3$ . For wavelengths larger than 800 nm, we suspect that the inversion routine did not distinguish well between  $a_1$  and  $a_3$  for cress, resulting in the negative values for  $a_1$ . However, the impact of  $a_1$  relative to  $a_0$  is negligible for cress for  $a_0 > 0.3$ .

Coefficient  $a_3$  (plots J-L) tends to zero for small  $a_0$  and reaches a maximum at about  $a_0 = 0.25$  for grass. We don't know whether this spectral behaviour is consistent with the assumption that  $a_3$  can be interpreted as a specular peak. If the Fresnel Reflectance of the scattering surface was constant (which is propably not the case here),  $a_3$  would be constant.  $a_3$  for cress is about a factor of 10 smaller than  $a_3$  for grass.

If there was no multiple scattering (MS), we would expect the relation of  $a_2$  (the hotspot intensity) and  $a_0$  to be linear. Obviously, this relation is not linear (plots G-I). The reason is, that as  $a_0$  increases, the intensity of MS increases too. MS reduces the hotspot because the shadowed components of the canopy are less dark since they are illuminated by the MS. Thus, for large values of  $a_0$  (corresponding to high albedos)  $a_2$  rises much slower than for small values of  $a_0$ .

For small  $a_0$  and  $a_2$ , (2) can be transformed to  $x \approx \frac{a_0}{a_0+a_2}$ . x was estimated from  $\frac{a_0}{a_0+a_2}$ , an exact determination was not possible due to the ambiguous behaviour of  $\frac{a_0}{a_0+a_2}$  for small  $a_0$ . k and c of (2) were fitted to all values of  $a_0$  and  $a_2$ . As can be seen from the solid lines in plots G-I, (2) describes the relation of  $a_0$  and  $a_2$  very well. Note that (2) covers values of  $a_0$  from 0.014 to 0.306 for grass (lab.) and 0.019 to 0.561 for cress, which is a difference of more than a factor of 20.

The coefficient k is proportional to the multiple scattered radiation. The difference of the coefficient k for grass and cress can be explained by the canopy structure: cress is a planophile canopy, not allowing much MS, whereas grass is erectophile with a lot of possibilities for MS. k for the laboratory measurement of grass is lower than k for the field measurement of grass. This is consistent with the above explanation, because the grass in the laboratory measurement has longer blades and is therefore not as erectophile as the field grass sample, and thus produces less MS. Coefficient c for the grass laboratory sample deviates significantly from 1, this may indicate problems with the application of (2) to this sample.

## CONCLUSIONS

We described the BRDF of grass and cress by an empirical function, consisting of a lambertian component, a hotspot, a bowlshape and a forward scattering term. We derived a physical relation between the intensity of the hotspot  $(a_2)$  and the lambertian component  $(a_0)$ . This relation is capable of describing the relation of  $a_0$  and  $a_2$  for grass and cress very well.

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*Fig.1:* Wavelength dependent coefficients (rhombs) of (1) for the 3 samples as a function of wavelength (plots A-C) and  $a_0$ , respectively (plots D-L). The solid lines in plots G-I denote relation (2) using k, c and x from table 1.

Sample	$b_1  [rad^{-4}]$	$b_2 [{\rm rad}^{-1}]$	$b_3  [rad^{-4}]$	$b_4  [rad^{-2}]$	k	С	Х	d (red)	d (NIR)
Grass (Lab.)	1.5	1.72	4.0	0.95	0.43	0.73	0.1	0.063	0.035
Grass (Field)	1.5	1.72	4.0	0.95	0.55	0.99	0.1	0.052	0.034
Cress	1.5	2.5	4.0	0.95	0.25	1.01	0.4	0.041	0.036

*Table 1: Wavelength independent coefficients and mean relative deviation* d  $(d = \frac{1}{N} \sum_{i=1,N} \frac{|BRF_i^{measured} - BRF_i^{modeled}|}{BRF_i^{measured}})$ .